

Flux $F \uparrow$: épu diff reliant à M^B

II.1 objectif : flux surfaciques de bande en IR technique

- Hyp :
- IR technique \Rightarrow source thermique $T_V = T_E (T_E)$
 - atmosphère plane // x et y indépendamment flux
 - symétrique / angle φ
 - pas diffusion $\omega_s = 0$
 - absorption forte \Rightarrow saturation $\Rightarrow \Delta z_V \approx 1$ pour interaction forte

Méthode

x solution formelle de l'équ du T.R.

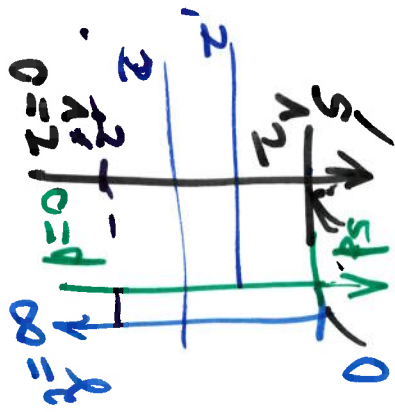
x réduite en de variables par $\int d\varphi$ $\int \mu d\mu$

x $\int_{\bar{v}-\frac{\Delta v}{2}}^{\bar{v}+\frac{\Delta v}{2}} d\bar{v}$

$\Delta v \gg$ largeur de raies

$\Delta \bar{v}$ faible pour B_r au voisinage pas bcp sur $\Delta \bar{v}$

II.2 Intégration angulaire



$$L^{\uparrow}(z_v, \mu, \varphi) = B_v(\tau_s) e^{-\tau_s - z_v} + \int_{z_v}^{\infty} \text{couches inf.} B_v(\tau_{z'}) e^{-\mu \frac{dz'}{\mu}}$$

$$L^{\downarrow}(z_v, \mu, \varphi) = 0 + \int_0^{z_v} \text{couches supér.} B_v(\tau_{z'}) e^{-\mu \frac{dz'}{\mu}}$$

$$F^{\uparrow}(z_v) = \int_{\varphi=0}^{2\pi} \int_{\theta=0}^{\pi/2} L^{\uparrow}(z_v, \mu, \varphi) \cos \theta \, d^2 \Omega \int_{\mu=0}^{\infty} \mu \sin \theta \, d\mu \, d\varphi$$

$$F^{\uparrow}(z_v) = 2\pi \int_{\mu=0}^{\infty} \int_{\theta=0}^{\pi/2} L^{\uparrow}(z_v, \mu) \mu \, d\mu \, d\theta$$

$$F^{\downarrow}(z_v) = 2\pi \int_{\mu=0}^{\infty} L^{\downarrow}(z_v, \mu) \mu \, d\mu$$

(2)

$T^*(z_v)$

$$\int_0^1 e^{-\frac{z_v - z_s}{\mu}} \mu dp$$

$$+ 2\pi \int_{z_v}^{z_s} B_v(T_{z_v}) \int_0^1 e^{-\frac{z_v - z_v}{\mu}} \mu dp$$

$$F^\downarrow(z_v) = 2\pi \int_{z_v}^{z_v} B_v(T_{z_v}) \int_0^1 e^{-\frac{z_v - z_v}{\mu}} \mu dp$$

II-3 Transmission diffuse

Transmissivité	directionnelle	diffuse
Rapport de luminosité	lumineuses	flux verticaux surfaciques
monochromatique	$T(z_v, \mu) = e^{-\frac{z_v}{\mu}}$	$T^*(z_v) = \int_0^1 \int_0^1 e^{-\frac{z_v}{\mu}} \mu dp$
de bande	$\langle T(z_v, \mu) \rangle$	$\int_0^1 \int_0^1 \mu dp$
$[\bar{\nu} - \frac{\Delta\nu}{2}, \bar{\nu} + \frac{\Delta\nu}{2}]$	$[\bar{\nu} - \frac{\Delta\nu}{2}, \bar{\nu} + \frac{\Delta\nu}{2}]$	$[\frac{\Delta\nu}{2}, \frac{\Delta\nu}{2}]$

avec μ positif dans l'axe z

Exponentialintegralfunktion E_n

$$E_n(z) = \int_1^{\infty} \frac{e^{-xz}}{x^n} dx$$

$$y = \frac{1}{x} \quad dy = -\frac{dx}{x^2}$$

$$dx = -\frac{dy}{y^2}$$

$$E_n(z) = \int_1^{\infty} e^{-\frac{z}{y}} y^{n-2} dy$$

$$\frac{dE_n}{dz} = - \int_1^{\infty} e^{-\frac{z}{y}} y^{n-3} dy = -E_{n-1}(z)$$

$$T^*(z) = 2 \int_0^1 e^{-\frac{z}{y}} y^2 dy = 2 E_3(z)$$

$$\frac{\partial T^*(z)}{\partial z} = -2 \int_0^1 e^{-\frac{z}{y}} y dy = -2 E_2(z) < 0$$

Expression de F^{\uparrow} en fonction de T^* et $\mathcal{D}T^*$

(15)

$$F^{\uparrow}(z_v) = M_v^{\beta} \left(T_{\text{surf}}^{\uparrow} \right) \int_{\text{surf}} T^* \left(z_v \right) \int_{z_v}^{\infty} M_v^{\beta} \left(T^* \right) \int_{z_v}^{\infty} \mathcal{D}T^* \left(z_v \right) = \int_{z_v}^{\infty} M_v^{\beta} \left(T^* \right) \int_{z_v}^{\infty} \mathcal{D}T^* \left(z_v \right)$$

$$F^{\uparrow}(z_v) = \int_{z_v}^{\infty} M_v^{\beta} \left(T^* \right) \int_{z_v}^{\infty} \mathcal{D}T^* \left(z_v \right)$$

II.4 Approx du noyau exponentielle

$$v \text{ fixe } T^*(z_v) = \int_0^1 e^{-z_v/p} \mu dp = \mathcal{E}_3(z_v) = e^{-\frac{z_v}{\bar{p}}} = T(z_v, \bar{p})$$

$\bar{p}_v =$ fonction de z_v , donc de v

Expression des flux en fait de l'épaisseur optique oblique (16)

$$z_{\bar{v}}^* = \frac{\tau_{\bar{v}}}{\bar{v}} \text{ oblique} = \left(\frac{z_{\bar{v}}^*}{s} \right) + \int_{z_{\bar{v}}^*}^{z_{\text{surf}}} n_{\bar{v}}(z_{\bar{v}}^*) e^{-\tau_{\bar{v}}^*} dz_{\bar{v}}^*$$

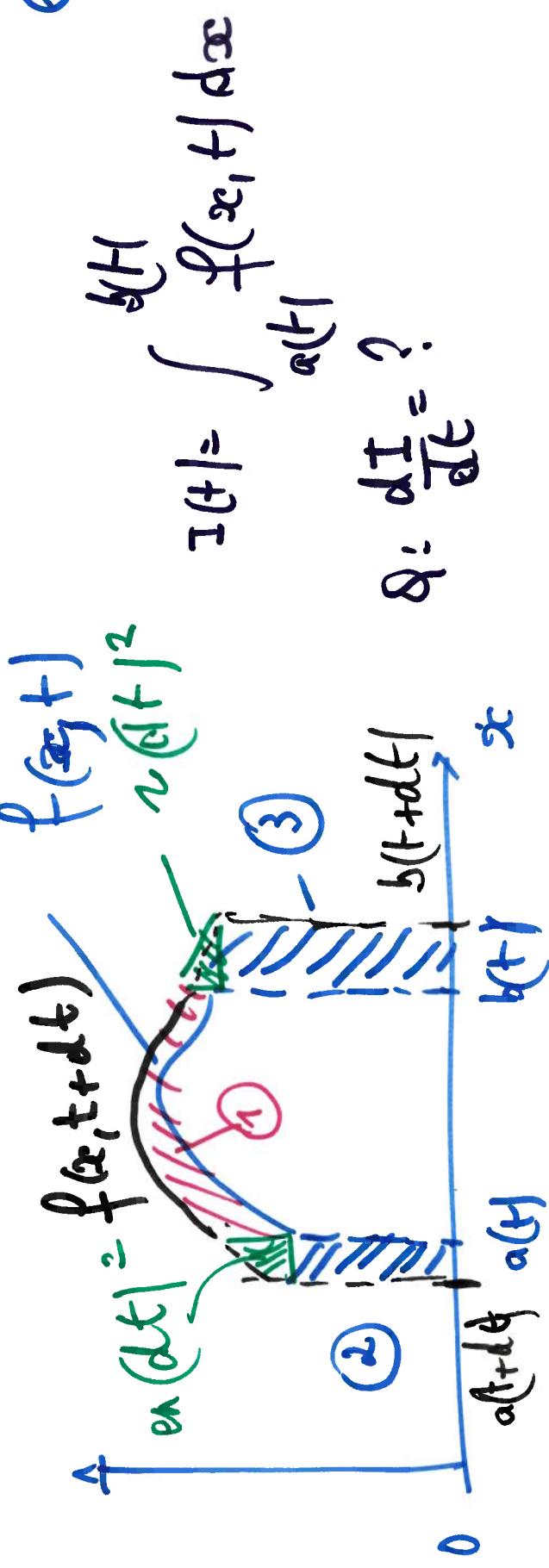
$$F_{\bar{v}}^{\uparrow}(z_{\bar{v}}^*) = \int_0^{z_{\bar{v}}^*} M_{\bar{v}}^B(z_{\bar{v}}^*) e^{-\tau_{\bar{v}}^*} dz_{\bar{v}}^*$$

$$F_{\bar{v}}^{\uparrow}(z) = M(z) e^{-\tau_s} + \int_{z_s}^{z_{\text{surf}}} M(z) e^{-\tau_s} dz$$

$$F_{\bar{v}}^{\downarrow}(z) = \int_0^z M(z) e^{-\tau_s} dz$$

$\frac{dF_{\bar{v}}^{\uparrow}}{dz}$
 $\frac{dF_{\bar{v}}^{\downarrow}}{dz}$

(17)



$$\frac{dI}{dt} = \frac{d}{dt} \int_{a(t)}^{b(t)} f(x, t) dx \quad (1)$$

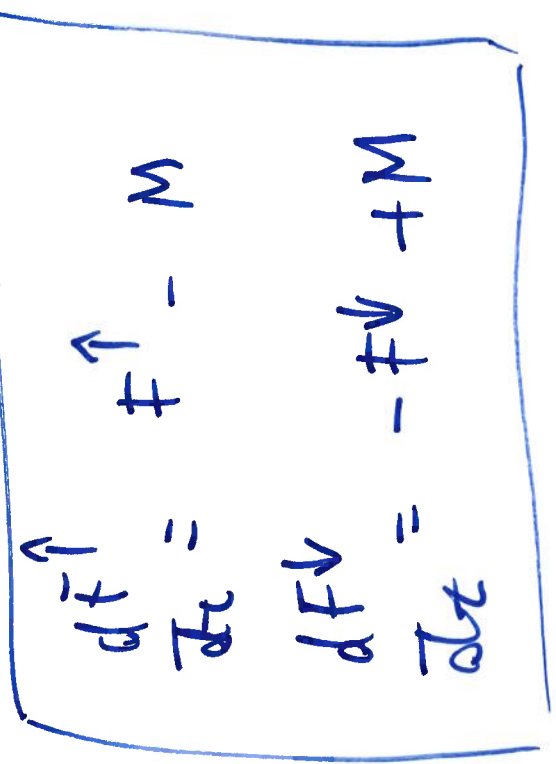
$$(2) \sim \frac{\partial a}{\partial t} f(a, t) + \int_{a(t)}^{b(t)} \frac{\partial f}{\partial t} dx + \int_{a(t+dt)}^{b(t+dt)} f(x, t) dx$$

$$(3) \sim \frac{\partial b}{\partial t} f(b, t) + \int_{a(t)}^{b(t)} f(x, t) dx$$

$$\frac{dI}{dt} = \frac{\partial a}{\partial t} f(a, t) + \int_{a(t)}^{b(t)} \frac{\partial f}{\partial t} dx + \frac{\partial b}{\partial t} f(b, t) + \int_{a(t)}^{b(t)} f(x, t) dx$$

$$\frac{dF^\uparrow}{dt} = \underbrace{M(s) e^{-(z-s_2)} + \int_{s_2}^z M(z) e^{-(z-z')} - M(z)}_{F^\uparrow - M}$$

$$\frac{dF^\downarrow}{dt} = M(z) - \int_0^z M(z) e^{-(z-z')} = M - F^\downarrow$$



III Méthodes à deux flux

III.1 Introduction à l'équilibre radiatif troposphérique

Hyp : - atmosphère transparente au solaire
- sans albedo (pas d'absorption)

- énergétique $\Rightarrow \int dV$:
atmosphère gazeuse diffuse et pure en IR thermique
 z_V indépendante de ν $F_{\uparrow}(z) = \int_{\nu} F_{\nu}^{\uparrow}(z_{\nu}) d\nu$

C. Limites

- $\Phi_0 = \frac{(1-A)C}{4}$ absorbé en surface
 $F_{\uparrow}(z=0)$ doit équilibrer Φ_0

- surface = C-N.

Supposant l'équilibre radiatif $F_{net} = F_{\uparrow} - F_{\downarrow}$ n'est de z
 $\frac{\partial F}{\partial z} = 0$

III.2 Résolution

$$F = F \uparrow - F \downarrow \quad \bar{F} = F \uparrow + F \downarrow$$

$$\left. \begin{array}{l} \frac{dF \uparrow}{dt} = F \uparrow - M \\ \frac{dF \downarrow}{dt} = -F \downarrow + M \end{array} \right| \begin{array}{l} \ominus \\ \oplus \end{array} \left| \begin{array}{l} \frac{d\bar{F}}{dt} = \bar{F} - 2M = 0 \\ \frac{d\bar{F}}{dt} = F = F(0) = \Phi_0 \end{array} \right.$$

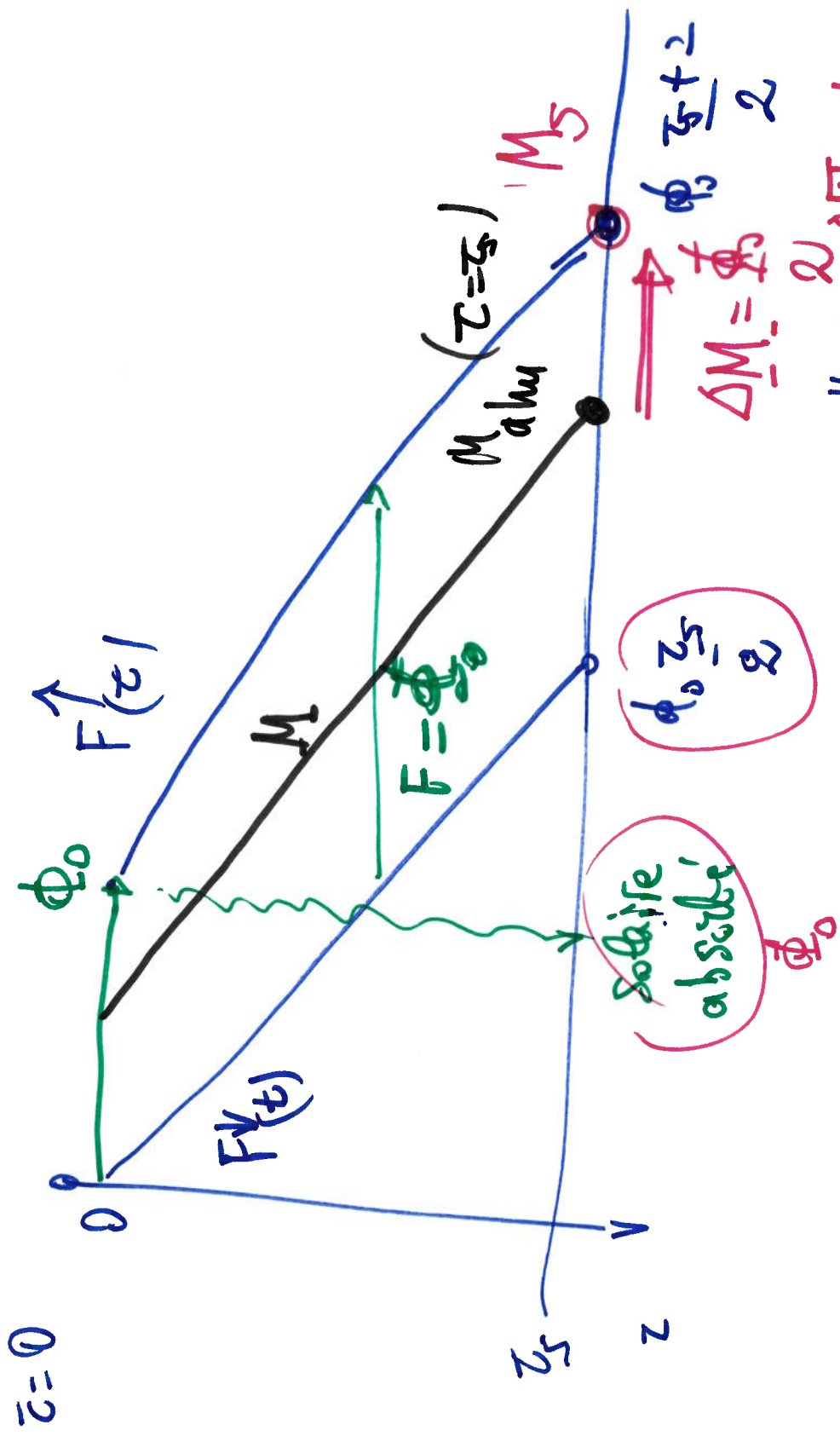
$$\bar{F} = \bar{F}(0) + \Phi_0 \tau \quad \bar{F}(0) = F \downarrow(0) + F \uparrow(0) = 0 + \Phi_0$$

$$\left\{ \begin{array}{l} \bar{F} = \Phi_0(1 + \tau) \\ F = \Phi_0 \end{array} \right. \left| \begin{array}{l} F \uparrow = \frac{\bar{F} + F}{2} \\ F \downarrow = \frac{\bar{F} - F}{2} \end{array} \right. \left. \begin{array}{l} F \uparrow = \Phi_0 \frac{\tau + 2}{2} \\ F \downarrow = \Phi_0 \frac{\tau}{2} \end{array} \right.$$

$$M = \Phi_0 \frac{\tau + 1}{2}$$

Solutions

(21)



$\Rightarrow \Delta T$ entre surf et atmos
 Equilibre radiatif impossible